# Markscheme 

May 2017

Physics

Higher level

Paper 3

22 pages

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## Section A

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | a |  | in order to keep the temperature constant $\checkmark$ in order to allow the system to reach thermal equilibrium with the surroundings/OWTTE $\checkmark$ | Accept answers in terms of pressure or volume changes only if clearly related to reaching thermal equilibrium with the surroundings. | 1 max |
|  | b |  | recognizes $b$ as gradient $\checkmark$ calculates $b$ in range $4.7 \times 10^{4}$ to $5.3 \times 10^{4} \checkmark$ Pam $\checkmark$ | Award [2 max] if POT error in $b$. <br> Allow any correct SI unit, eg $\mathrm{kg} \mathrm{s}^{-2}$. | 3 |
|  | C |  | $V \propto H$ thus ideal gas law gives $p \propto \frac{1}{H} \checkmark$ <br> so graph should be «a straight line through origin,» as observed $\checkmark$ |  | 2 |
|  | d |  | $\begin{aligned} & n=\frac{b A}{R T} O R \text { correct substitution of one point from the graph } \\ & n=\frac{5 \times 10^{4} \times 1.3 \times 10^{-3}}{8.31 \times 300}=0.026 \approx 0.03 \end{aligned}$ | Answer must be to 1 or 2 SF. Award [2] for a bald correct answer. <br> Allow ECF from (b). | 2 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | e |  | very large $\frac{1}{H}$ means very small volumes / very high pressures at very small volumes the ideal gas does not apply <br> OR <br> at very small volumes some of the assumptions of the kinetic theory of gases do not hold $\checkmark$ |  | 2 |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 2 | a | $\begin{aligned} & g=\frac{4 \pi^{2} \times 1.60}{2.540^{2}}=9.7907 \\ & \Delta g=g\left(\frac{\Delta L}{L}+2 \times \frac{\Delta T}{T}\right)=« 9.7907 \times\left(\frac{0.01}{1.60}+2 \times \frac{0.005}{2.540}\right)=» 0.0997 \\ & \text { OR } \\ & 1.0 \% \checkmark \\ & \text { hence } g=(9.8 \pm 0.1) \text { «} \mathrm{ms}^{-2} » \text { OR } \Delta g=0.1 \text { « } \mathrm{ms}^{-2} » \checkmark \end{aligned}$ | For the first marking point answer must be given to at least 2 dp . <br> Accept calculations based on $\begin{aligned} & g_{\max }=9.8908 \\ & g_{\min }=9.6913 \\ & \frac{g_{\max }-g_{\min }}{2}=0.099 \approx 0.1 \end{aligned} .$ | 3 |
|  | b | $\begin{aligned} & \frac{T}{T_{0}}=1.01 \checkmark \\ & \theta_{\max }=22 «<» \end{aligned}$ | Accept answer from interval 20 to 24. | 2 |

## Section B

Option A - Relativity

| Question |  | Answers | Notes | Total |
| :---: | :---: | :--- | :--- | :--- | :---: |
| $\mathbf{3}$ | a | a set of coordinate axes and clocks used to measure the position «in space/time of an object <br> at a particular time» <br> OR <br> a coordinate system to measure x,y,z,and $t /$ OWTTE $\checkmark$ | $\mathbf{1}$ |  |
|  | b | i | magnetic only $\checkmark$ <br> there is a current but no «net» charge «in the wire» $\checkmark$ |  |
|  | b | ii | electric only $\checkmark$ <br> $P$ is stationary so experiences no magnetic force $\checkmark$ <br> relativistic contraction will increase the density of protons in the wire $\checkmark$ | $\mathbf{2}$ |



| d | ii |  |  |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 5 | a | $\begin{aligned} & \gamma=1.96 \checkmark \\ & E_{k}=(\gamma-1) m_{0} c^{2}=900 « M e \mathrm{~V} » \\ & p d \approx 900 \text { «MV» } \checkmark \end{aligned}$ | Award [2 max] if Energy and Potential difference are not clearly distinguished, eg by the unit. | 3 |
|  | b | ```energy of proton = }\mp@subsup{m}{mc}{2}=1838<MeV» total energy available = energy of proton + energy of antiproton = 1838+1838=3676 <MeV»\checkmark momentum of a one photon = Total energy / 2c = 1838 «Me Vc-1» \checkmark``` |  | 3 |


| $\mathbf{6}$ | $\mathbf{a}$ |  | $f=« \frac{E}{h}=» \frac{14400 \times 1.6 \times 10^{-19}}{6.63 \times 10^{-34}}=« 3.475 \times 10^{18} \mathrm{~Hz} » \checkmark$ <br> $\Delta f=« \frac{g \times \Delta h \times f}{c^{2}} \approx » 8550 《 \mathrm{~Hz} » \checkmark$ |  |
| :--- | :--- | :--- | :--- | :--- | :--- |
|  | $\mathbf{b}$ | «as the photon moves away from the Earth, » it has to spend energy to overcome the <br> gravitational field $\checkmark$ <br> since $E=h f$, the detected frequency would be lower «than the emitted frequency» $\checkmark$ | $\mathbf{2}$ |  |

## Option B — Engineering physics

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 7 | a | i | $\frac{M}{3} v R \checkmark$ |  | 1 |
|  | a | ii | evidence of use of: $L=I \omega=\left(M R^{2}+\frac{M}{3} R^{2}\right) \omega \checkmark$ |  | 1 |
|  | a | iii | evidence of use of conservation of angular momentum, $\frac{M v R}{3}=\frac{4}{3} M R^{2} \omega$ «rearranging to get $\omega=\frac{v}{4 R}$ " |  | 1 |
|  | a | iv | initial $K E=\frac{M v^{2}}{6} \checkmark$ <br> final $K E=\frac{M v^{2}}{24}$ <br> energy loss $=\frac{M v^{2}}{8}$ |  | 3 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 7 | b | i | $\begin{aligned} & \alpha<=\frac{3}{4} \frac{\Gamma}{M R^{2}} »=\frac{3}{4} \frac{0.01}{0.7 \times 0.5^{2}} \\ & \text { «to give } \alpha=0.04286 \mathrm{rads}^{-2} \text { " } \end{aligned}$ | Working OR answer to at least 3 SF must be shown. | 1 |
|  | b | ii | $\begin{aligned} & \theta=\frac{\omega_{i}^{2}}{2 \alpha} « \text { from } \omega_{t}^{2}=\omega_{i}^{2}+2 \alpha \theta » \checkmark \\ & \theta «=\frac{v^{2}}{32 R^{2} \alpha}=\frac{2.1^{2}}{32 \times 0.5^{2} \times 0.043} »=12.8 \text { OR } 12.9 \text { «rad» } \\ & \text { number of rotations «=} \frac{12.9}{2 \pi} »=2.0 \text { revolutions } \checkmark \end{aligned}$ |  | 3 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 8 | a |  | «a process in which there is» no thermal energy transferred between the system and the surroundings $\checkmark$ |  | 1 |
|  | b |  | A to $\mathrm{B} \boldsymbol{A N D} \mathrm{C}$ to $\mathrm{D} \checkmark$ |  | 1 |
|  | c | i | $\begin{aligned} & T=\frac{P V}{n R} \checkmark \\ & T\left(=\frac{512 \times 10^{3} \times 1.20 \times 10^{-3}}{0.150 \times 8.31}\right) \approx 493 \text { «K» } \end{aligned}$ | The first mark is for rearranging. | 2 |
|  | c | ii | $\begin{aligned} & P_{B}=\frac{P_{a} V_{A}}{V_{B}} \checkmark \\ & P_{B}=267 \mathrm{kPa} \checkmark \end{aligned}$ | The first mark is for rearranging. | 2 |
|  | d | i | «B to C adiabatic so» $P_{B} V_{B}^{\frac{5}{3}}=P_{C} V_{C}^{\frac{5}{3}}$ AND $P_{C} V_{C}=n R T_{C}$ «combining to get result» $\checkmark$ | It is essential to see these 2 relations to award the mark. | 1 |
|  | d | ii | $\begin{aligned} & T_{C}=\left(\frac{P_{B} V_{B}^{\frac{5}{3}}}{n R}\right) V_{C}^{\frac{-2}{3}} \checkmark \\ & T_{C}=«\left(\frac{267 \times 10^{3} \times\left(2.30 \times 10^{-3}\right)^{\frac{5}{3}}}{0.150 \times 8.31}\right)\left(2.90 \times 10^{-3}\right)^{\frac{-2}{3}} 》=422 « \mathrm{~K} » \checkmark \end{aligned}$ |  | 2 |
|  | e |  | the isothermal processes would have to be conducted very slowly / OWTTE $\checkmark$ |  | 1 |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 9 | a | ALTERNATIVE 1 <br> pressure in a liquid increases with depth $\checkmark$ <br> so pressure at bottom of bubble greater than pressure at top $\checkmark$ <br> ALTERNATIVE 2 <br> weight of liquid displaced $\checkmark$ <br> greater than weight of bubble |  | 2 |
|  | b | $\frac{\text { weight }}{\text { buoyancy }}\left(=\frac{V \rho_{a} g}{V \rho_{l} g}=\frac{\rho_{a}}{\rho_{l}}=\frac{1.2}{1200}\right)=10^{-3}$ <br> since the ratio is very small, the weight can be neglected $\checkmark$ | Award [1 max] if only mass of the bubble is calculated and identified as negligible to mass of liquid displaced. | 2 |
|  | c | evidence of equating the buoyancy and the viscous force " $\rho_{l} \frac{4}{3} \pi r^{3} g=6 \pi \eta r v_{t}$ " $\checkmark$ $v_{t}=<\frac{2}{9} \frac{1200 \times 9.81}{1 \times 10^{-3}}\left(0.25 \times 10^{-3}\right)^{2}=» 0.16 《 \mathrm{~m} \mathrm{~s}^{-1} » \checkmark$ |  | 2 |


| $\mathbf{1 0}$ | $\mathbf{a}$ | the loss of energy in an oscillating system $\checkmark$ | $\mathbf{1}$ |  |
| :--- | :--- | :--- | :--- | :--- |
|  | $\mathbf{b}$ | $Q=2 \pi \frac{16^{2}}{16^{2}-10.3^{2}} \approx 11 \checkmark$ <br> the amplitude decreases at a slower rate $\checkmark$ <br> a higher $Q$ factor would mean that less energy is lost per cycle $\checkmark$ | Accept calculation based on <br> any two correct values giving <br> answer from interval 10 to 13. | $\mathbf{1}$ |
| $\mathbf{c}$ |  | $\mathbf{2}$ |  |  |

## Option C - Imaging

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 11 | a | i | an image formed by extensions of rays, not rays themselves OR <br> an image that cannot be projected on a screen $\checkmark$ |  | 1 |
|  | a | ii | $\begin{aligned} & \frac{1}{v}=\frac{1}{3.0}-\frac{1}{4.0} \checkmark \\ & « v=12 \mathrm{~cm} » \end{aligned}$ |  | 1 |
|  | a | iii | $\begin{aligned} & u=18-12=6.0 « \mathrm{~cm} » \\ & v=-24<\mathrm{cm} » \\ & \text { « } \frac{1}{f}=\frac{1}{6.0}-\frac{1}{24} \Rightarrow » f=8.0<\mathrm{cm} » \end{aligned}$ | Award [2 max] for answer of 4.8 cm . <br> Minus sign required for MP2. | 3 |
|  | a | iv | line parallel to principal axis from intermediate image meeting eyepiece lens at $P \checkmark$ line from arrow of final image to $P$ intersecting principal axis at $F$ |  | 2 |


| Question |  | Answers | Notes |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- |
| $\mathbf{1 1}$ | $\mathbf{b}$ | $\mathbf{i}$ | object is far away so intermediate image forms at focal plane of objective $\checkmark$ <br> for final image at infinity object must also be at focal point of eyepiece $\checkmark$ <br> «hence 87.5 cm» | No mark for simple addition of <br> focal lengths without <br> explanation. |  |
|  | $\mathbf{b}$ | $\mathbf{i i}$ | angular magnification $=\frac{85.0}{2.50}=34 \checkmark$ <br> angular diameter $34 \times 7.8 \times 10^{-3}=0.2652 \approx 0.27$ «rad» $\checkmark$ <br> chromatic aberration is the dependence of refractive index on wavelength $\checkmark$ <br> but mirrors rely on reflection <br> OR <br> mirrors do not involve refraction $\checkmark$ <br> «so do not suffer chromatic aberration» | $\mathbf{2}$ |  |
| c | $\mathbf{2}$ |  |  |  |  |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 12 | a | i | longer distance without amplification signal cannot easily be interfered with $\checkmark$ less noise $\checkmark$ no cross talk $\checkmark$ higher data transfer rate $\checkmark$ |  | 2 max |
|  | a | ii | infrared radiation suffers lower attenuation $\checkmark$ |  | 1 |
|  | b |  | $\begin{aligned} & \text { loss }=10 \log \frac{2.4}{15} «=-7.959 \mathrm{~dB} » \\ & \text { length }=« \frac{7.959}{0.30}=» 26.53 \approx 27 \text { «km» } \downarrow \end{aligned}$ |  | 2 |
|  | c |  | a thin core means that rays follow essentially the same path / OWTTE $\checkmark$ and so waveguide (modal) dispersion is minimal / OWTTE $\checkmark$ |  | 2 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 13 | a |  | bone and tissue absorb different amounts of X-rays <br> OR <br> bone and tissue have different attenuation coefficients $\checkmark$ <br> so boundaries and fractures are delineated in an image $\checkmark$ |  | 2 |
|  | b |  | $\begin{aligned} & \frac{I_{\text {bone }}}{I_{\text {tissue }}}=\frac{I_{0} e^{-\mu_{b} x}}{I_{0} e^{-\mu_{4} x}}=e^{-\left(\mu_{b}-\mu_{t}\right) x} \\ & \frac{I_{\text {bone }}}{I_{\text {tissue }}}=e^{-1.2 \times 10^{-2} \times(1.9-1.1) \times 10^{3} \times 5.4 \times 10^{-2}} \\ & \frac{I_{\text {bone }}}{I_{\text {tissue }}}=0.60 \checkmark \end{aligned}$ |  | 3 |
|  | c | i | to split the energy level of protons in the body <br> OR <br> to cause protons in the body to align with the field / precess at Larmor frequency $\checkmark$ |  | 1 |
|  | C | ii | to force/excite protons that are in the spin up/parallel state $\checkmark$ into a transition to the spin down/antiparallel state $\checkmark$ |  | 2 |
|  | C | iii | the emitted radio frequency signal has a frequency that depends on the magnetic field $\checkmark$ with a gradient field different parts of the body have different frequencies and so can be identified |  | 2 |

## Option D - Astrophysics

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 14 | a | i | stars fusing hydrogen «into helium» $\checkmark$ |  | 1 |
|  | a | ii | $\begin{aligned} & M=M_{\odot}\left(4 \times 10^{5}\right)^{\frac{1}{3.5}}=39.86 M_{\odot} \\ & « M \approx 40 M_{\odot} " \end{aligned}$ | Accept reverse working. | 1 |
|  | a | iii | $\begin{aligned} & 4 \times 10^{5}=13^{2} \times \frac{T^{4}}{6000^{4}} \\ & T \approx 42000 « \mathrm{~K} » \end{aligned}$ | Accept use of substituted values into $L=\sigma 4 \pi R^{2} T^{4}$. | 2 |
|  | a | iv | $\begin{aligned} & 4 \times 10^{-11}=4 \times 10^{5} \times \frac{1 A U^{2}}{d^{2}} \\ & d=1 \times 10^{8} « A U » \checkmark \end{aligned}$ | Accept use of correct values into $b=\frac{L}{4 \pi d^{2}}$. | 2 |
|  | b |  | the gravitation «pressure» is balanced by radiation «pressure» $\checkmark$ that is created by the production of energy due to fusion in the core / OWTTE $\checkmark$ | Award [1 max] if pressure and force is inappropriately mixed in the answer. <br> Award [1 max] for unexplained "hydrostatic equilibrium is reached". | 2 |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 14 | C | the Sun will evolve to become a red giant whereas Theta 1 Orionis will become a red super giant $\checkmark$ <br> the Sun will explode as a planetary nebula whereas Theta 1 Orionis will explode as a supernova <br> the Sun will end up as a white dwarf whereas Theta 1 Orionis as a neutron star/black hole $\checkmark$ |  | 3 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 15 | a | i | black body radiation / $3 \mathrm{~K} \checkmark$ <br> highly isotropic / uniform throughout <br> OR <br> filling the universe | Do not accept: CMB provides evidence for the Big Bang model. | 2 |
|  | a | ii | $« \lambda=\frac{2.9 \times 10^{-3}}{2.8} \geqslant \approx 1.0 \ll \mathrm{~mm} » \checkmark$ |  | 1 |
|  | b |  | the universe is expanding and so the wavelength of the CMB in the past was much smaller $\checkmark$ indicating a very high temperature at the beginning $\checkmark$ |  | 2 |
|  | c | i | $\begin{aligned} & « z=\frac{v}{c} \Rightarrow » v=0.16 \times 3 \times 10^{5} «=0.48 \times 10^{5} \mathrm{~km} \mathrm{~s}^{-1} » \checkmark \\ & « d=\frac{v}{H_{0}} \Rightarrow v=\frac{0.48 \times 10^{5}}{68}=706 » \approx 710 « \mathrm{Mpc} » \end{aligned}$ | Award [1 max] for POT error. | 2 |
|  | c | ii | $\begin{aligned} & z=\frac{R}{R_{0}}-1 \Rightarrow \frac{R}{R_{0}}=1.16 \checkmark \\ & \frac{R_{0}}{R}=0.86 \checkmark \end{aligned}$ |  | 2 |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 16 | a | a star will form out of a cloud of gas $\checkmark$ <br> when the gravitational potential energy of the cloud exceeds the total random kinetic energy of the particles of the cloud <br> OR <br> the mass exceeds a critical mass for a particular radius and temperature $\checkmark$ |  | 2 |
|  | b | number of reactions is $\frac{10^{10} \times 365 \times 24 \times 3600 \times 3.8 \times 10^{26}}{4.3 \times 10^{-12}}=2.79 \times 10^{55} \checkmark$ H mass used is $2.79 \times 10^{55} \times 4 \times 1.67 \times 10^{-27}=1.86 \times 10^{29}$ «kg» $\checkmark$ |  | 2 |
|  | c | nuclear fusion reactions produce ever heavier elements depending on the mass of the star / temperature of the core <br> the elements / nuclear reactions arrange themselves in layers, heaviest at the core lightest in the envelope $\checkmark$ |  | 2 |


| Question |  | Answers | Total |  |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| $\mathbf{1 7}$ | $\mathbf{a}$ | curve starting earlier, touching at now and going off to infinity $\checkmark$ |  | Notes |

